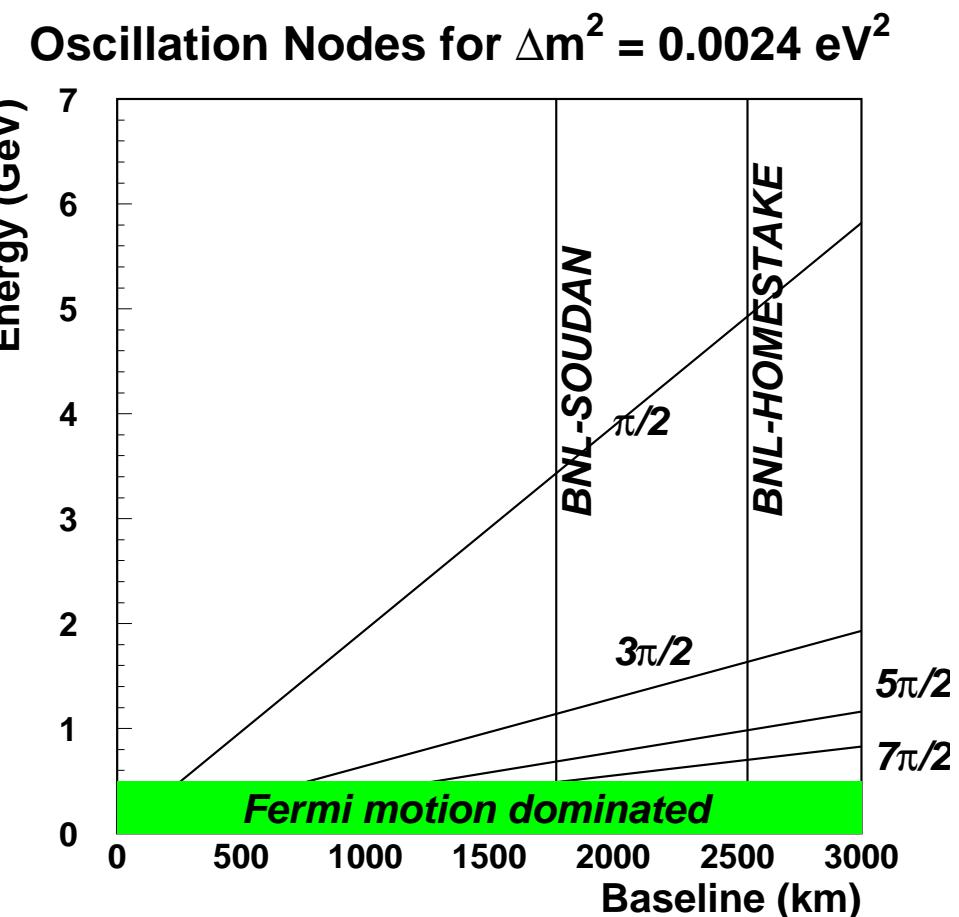


Very Long Baseline Scenarios

Milind Diwan March 2, 2004 Superbeam study

- A next generation large detector at a national underground facility with a powerful neutrino beam and very long baseline.
- Unique, and broad range of physics.
- Ability to over-constrain the neutrino mixing sector.
(3-generation mixing: δ_{CP} can be measured without running both polarities)
- This talk on Flexibility:
 - Not affected if parameters change.
 - Go off-axis for better detector performance.
 - Change running strategy if physics emphasis changes.

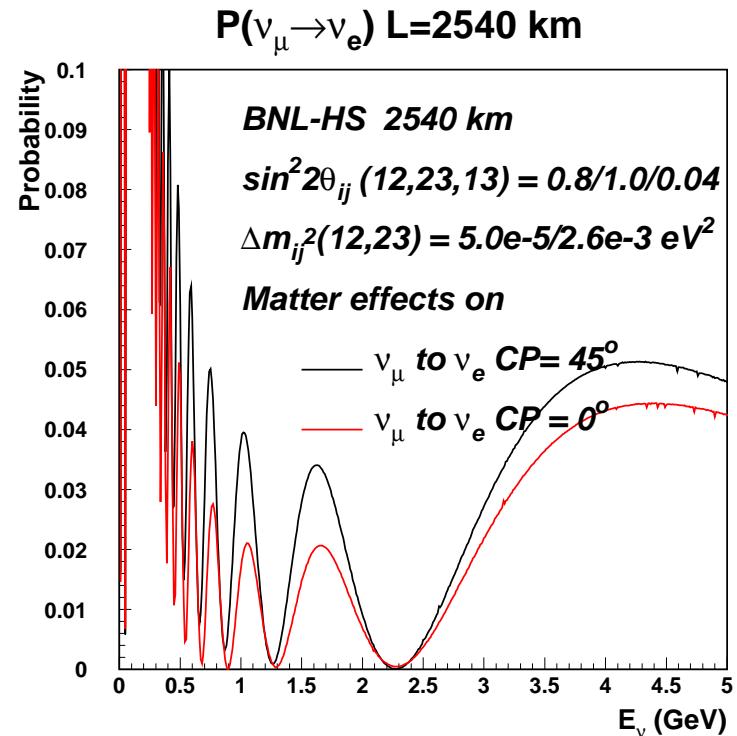
- Large effects: Multiple oscillation nodes.
- Low cross section at low energies
- Fermi motion limits resolution at low energies:
- $\Delta m^2 \approx 0.0024 \text{ eV}^2$: Baseline $\sim 2000 \text{ km}$.



BNL strategy for leptonic CP violation

General Features

- 0.5–1 GeV: Δm_{21}^2 (LMA) region.
- 1 – 3 GeV: CP large effects region
- > 3 GeV: Matter enhanced (ν_μ), suppressed ($\bar{\nu}_\mu$). ($\Delta m_{32}^2 > 0$).



Using neutrinos alone make fit to spectrum to extract δ_{CP} .

Full implementation: wide band beam over very long distance.

$\nu_\mu \rightarrow \nu_e$ with matter effect

Approximate formula (Lindener, Huber et al.)

$$\begin{aligned}
P(\nu_\mu \rightarrow \nu_e) \approx & \sin^2 \theta_{23} \frac{\sin^2 2\theta_{13}}{(\hat{A} - 1)^2} \sin^2((\hat{A} - 1)\Delta) \\
& + \alpha \frac{8J_{CP}}{\hat{A}(1 - \hat{A})} \sin(\Delta) \sin(\hat{A}\Delta) \sin((1 - \hat{A})\Delta) \\
& + \alpha \frac{8I_{CP}}{\hat{A}(1 - \hat{A})} \sin(\Delta) \cos(\hat{A}\Delta) \sin((1 - \hat{A})\Delta) \\
& + \alpha^2 \frac{\cos^2 \theta_{23} \sin^2 2\theta_{12}}{\hat{A}^2} \sin^2(\hat{A}\Delta)
\end{aligned} \tag{1}$$

$$J_{CP} = 1/8 \sin \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$$

$$I_{CP} = 1/8 \cos \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$$

$$\alpha = \Delta m_{21}^2 / \Delta m_{31}^2, \Delta = \Delta m_{31}^2 L / 4E$$

$$\hat{A} = 2VE / \Delta m_{31}^2$$

Resolving Parameters with $\nu_\mu \rightarrow \nu_e$

Assume $L > 2000\text{km}$, wide band beam

Δm_{32}^2 , Δm_{21}^2 , θ_{12} well known.

3 neutrino generations. \uparrow = large change \uparrow = small change

		$\sin^2 2\theta_{13} > 0$	$\Delta m_{32}^2 (> 0, < 0)$	$\delta_{CP} = (\pi/4, -\pi/4)$	$\theta_{23} (< \pi/4, > \pi/4)$
ν	0 – 1.2 GeV	\uparrow	–, –	\uparrow, \downarrow	$\uparrow\downarrow, \downarrow\downarrow$
	1.2 – 2.2 GeV	\uparrow	–, –	$\uparrow\downarrow, \downarrow\downarrow$	$\downarrow\downarrow, \uparrow\uparrow$
	> 2.2 GeV	\uparrow	$\uparrow\downarrow, \downarrow\uparrow$	\uparrow, \downarrow	$\downarrow\downarrow, \uparrow\uparrow$
$\bar{\nu}$	0 – 1.2 GeV	\uparrow	–, –	\downarrow, \uparrow	$\uparrow\downarrow, \downarrow\downarrow$
	1.2 – 2.2 GeV	\uparrow	–, –	$\downarrow\downarrow, \uparrow\uparrow$	$\downarrow\downarrow, \uparrow\uparrow$
	> 2.2 GeV	\uparrow	$\downarrow\downarrow, \uparrow\uparrow$	\downarrow, \uparrow	$\downarrow\downarrow, \uparrow\uparrow$

Scaling Laws for CP Measurement

Effect of δ_{CP} compared to first term in appearance.

$R_{CP} \equiv$ Second term divided by First Term.

$$R_{CP} \propto \sin \delta_{CP} \frac{\Delta m_{21}^2 L}{4E} \frac{1}{\sin 2\theta_{13}}$$

- $R_{CP} \propto 1/E$. Matter effect only at high E .
Allows separation of matter effect and CP effect.
- $R_{CP} \propto L$. Event rate $\propto 1/L^2$.
Statistical merit indep. of L for same sized detector.
- $R_{CP} \propto 1/\sin 2\theta_{13}$. Electron event rate $\propto \sin^2 2\theta_{13}$.
Statistical merit indep. of θ_{13} .
- $R_{CP} \propto \Delta m_{21}^2$. Better CP resolution for higher Δm_{21}^2 .
- For given resolution on δ_{CP} detector size is independent of L .

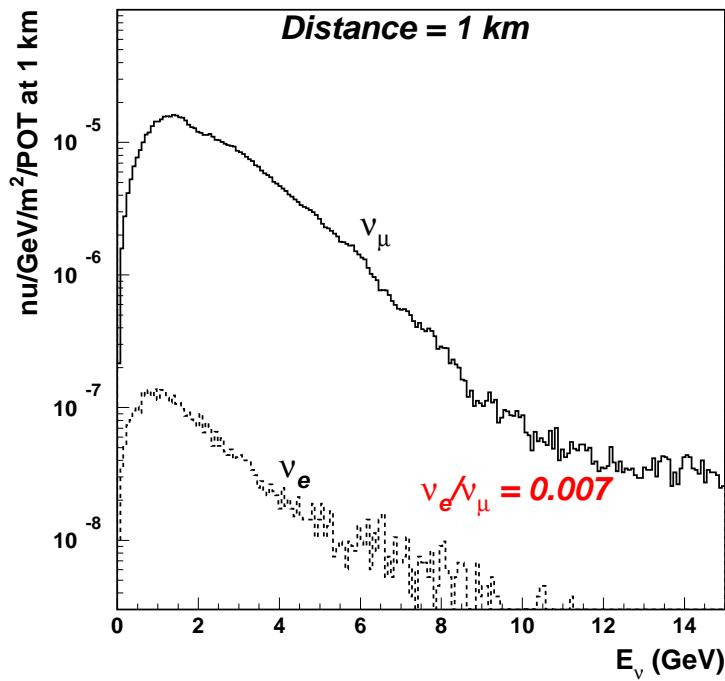
Running schemes and event rates

1 MW, 5×10^7 sec, Neutrino Running.

- Standard: Wideband beam to 500 kT detector at Homestake (2540 km) with strong background rejection.
CC: 52000 NC: 18000 QE: 11800
- Weak Backg. rej.: Wideband beam to 500 kT detector at Homestake (2540 km) with weak background rejection.
- Offaxis: 1 deg. Offaxis beam to 500 kT detector at Homestake (2540 km) with weak background rejection.
CC: 19000 NC: 7100 QE: 6500
- Some comments on 50 kT Liquid Argon detector.

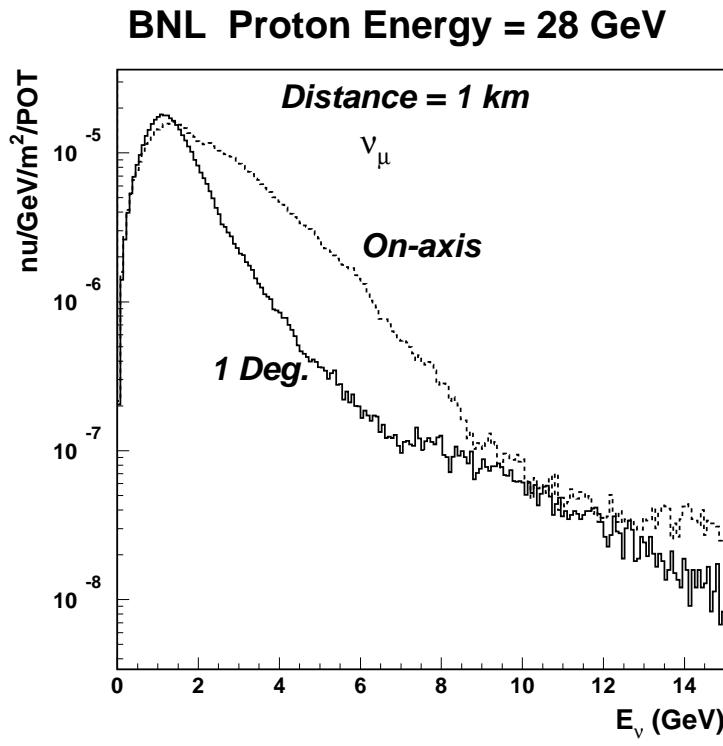
BNL on-axis

BNL Wide Band. Proton Energy = 28 GeV



- Flux and contamination experimentally known
- 60 cm graphite target.
- 4 m diameter, 200 m long tunnel
- $5 \times 10^{-5} \nu/\text{m}^2/\text{POT}$ @ 1 km.
- 52000 CC, 17000 NC events
(1MW, 2540 km, 0.5 MT, 5×10^7 sec)

1 deg. Off axis beam



- Move target and horn by 1.3 m and rotate 1 deg.
- Same 4 m diameter, 200 m long tunnel
- Will need large beam dump
- 19000 CC, 7000 NC events (1MW, 2540 km, 0.5 MT, 5×10^7 sec)

Scenario No. 1 Wide band

1 MW, Neutrinos, 500 kT, L=2540km, $5 \times 10^7 sec$

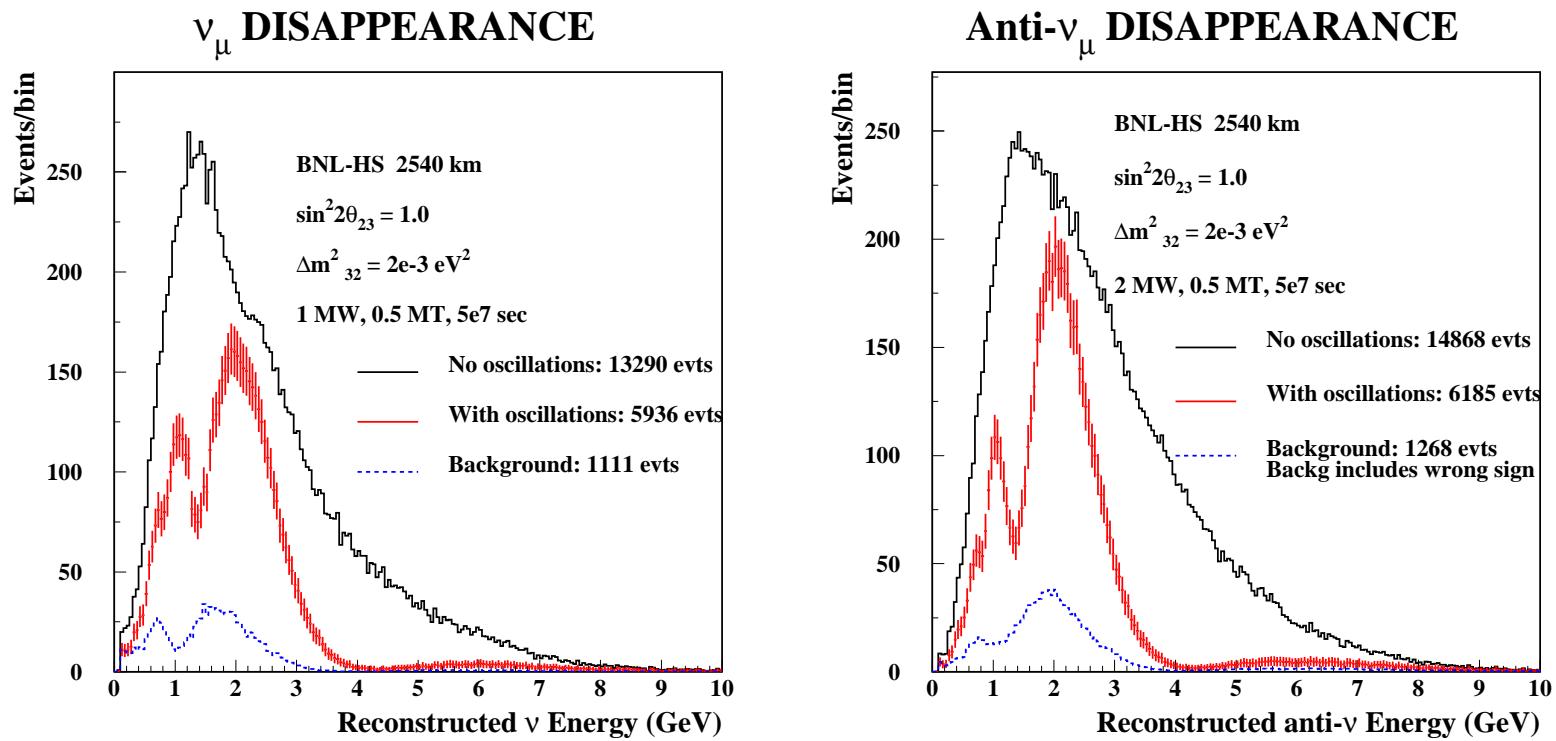
2 MW, Anti-Neutrinos, 500 kT, L=2540km, $5 \times 10^7 sec$

Strong Background Rejection.

CC $\nu_\mu N \rightarrow \mu^- X$	51800 (30050)	NC $\nu_\mu N \rightarrow \nu_\mu X$	18323 (11540)
CC $\nu_e N \rightarrow e^- X$	380 (106)		
QE $\nu_\mu n \rightarrow \mu^- p$	11767 (11868)	NC elastic	4575 (3882)
QE $\nu_e n \rightarrow e^- p$	84 (80)		
CC Single π	22053 (11872)	NC Single π	7741 (5074)
CC Two π	10143 (3336)	NC Two π	3557 (1630)
CC $> 2 \pi$	4882 (500)	NC $> 2 \pi$	1729 (560)
CC $\nu_\tau N \rightarrow \tau^- X$	~ 110 (40)	(depends on Δm^2)	

Anti-neutrino rate (brackets) for 2 MW.

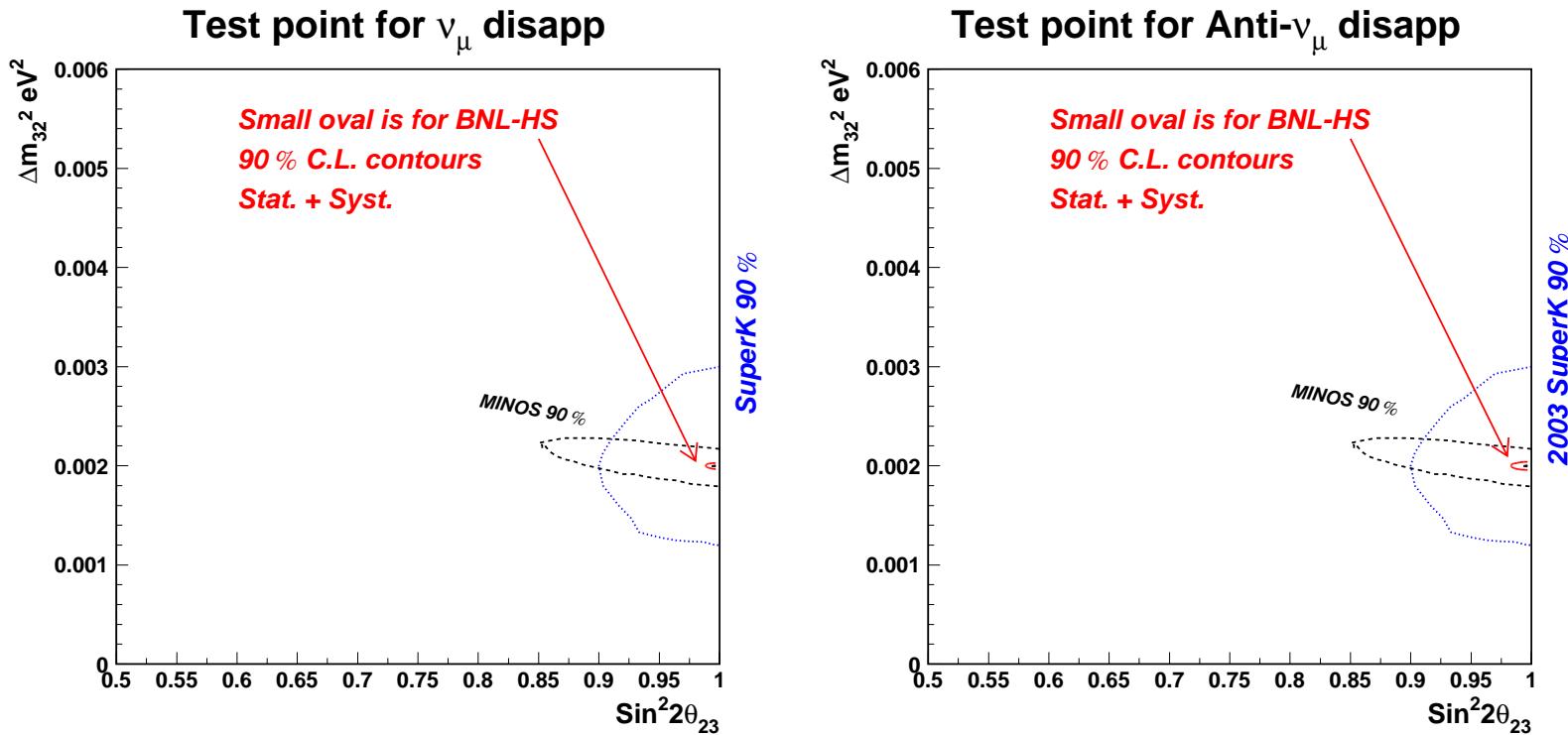
(Add Nu cont. to antinu rate: 7696 CC and 2600 NC.)



Node pattern provides high Δm^2_{32} resolution. Energy calibration is very important.

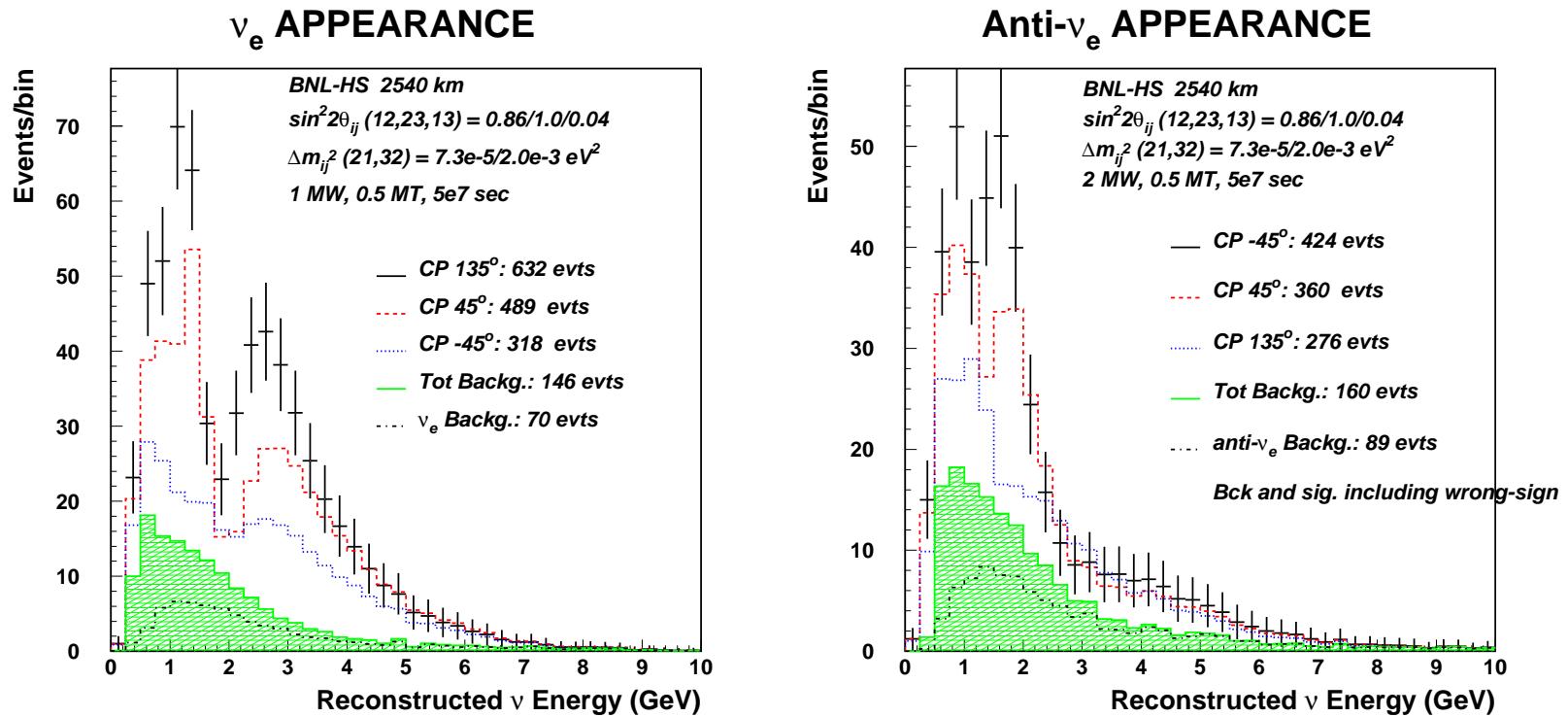
Flux normalization not important for measurement of $\sin^2 2\theta_{23}$
Minimum systematics in ν_μ and $\bar{\nu}_\mu$ comparison

Disappearance resolution, essential measurement of θ_{23}



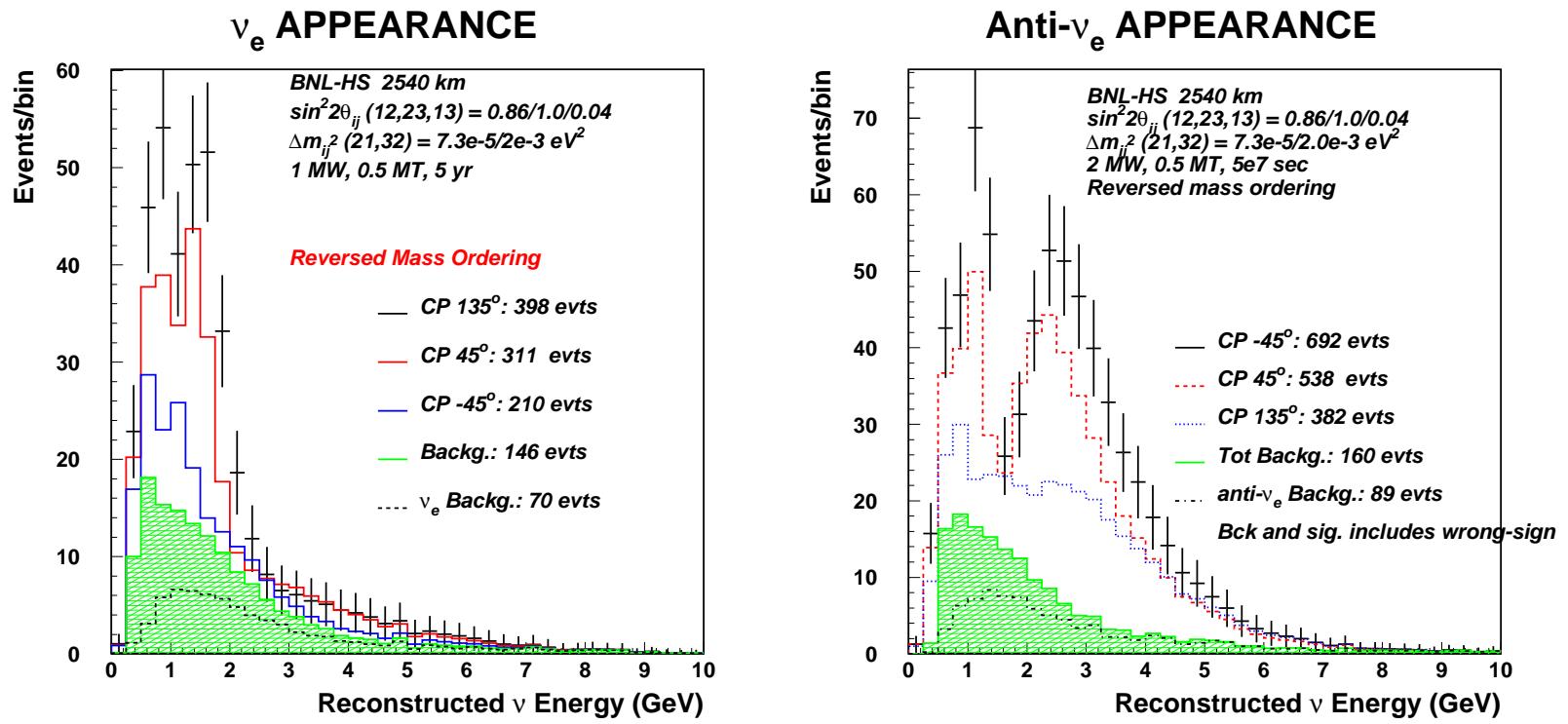
$\sim 1\%$ resol. on Δm_{32}^2 and $\sin^2 2\theta_{23}$ over broad range. Understand detector energy scale to 1%. Robust against systematics.

Running for $\sin^2 2\theta_{13}$



$\Delta m_{32}^2 = 0.002 \text{ eV}^2$, $\sin^2 2\theta_{13} = 0.04$. Assume normal mass hierarchy. $m_3 > m_2 > m_1$ Matter effects included.

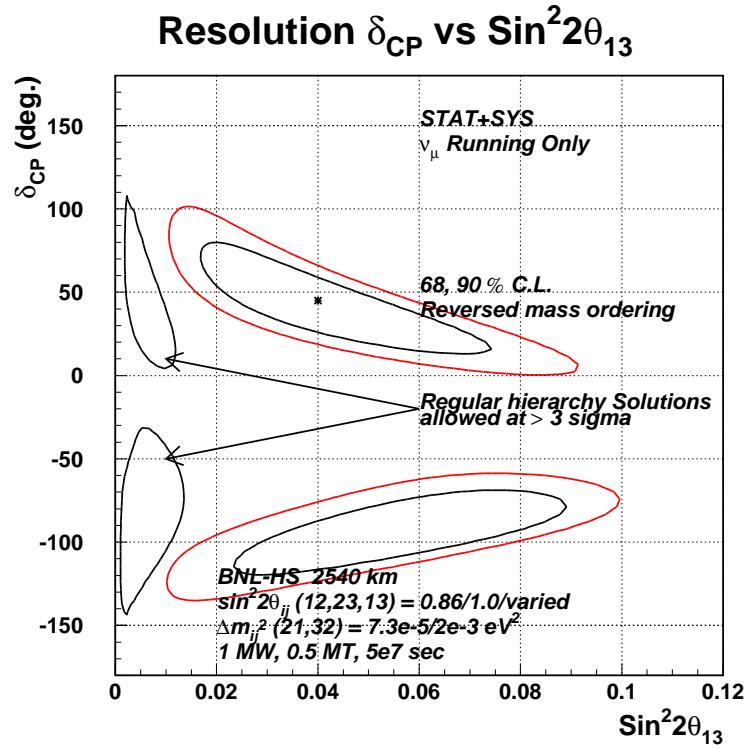
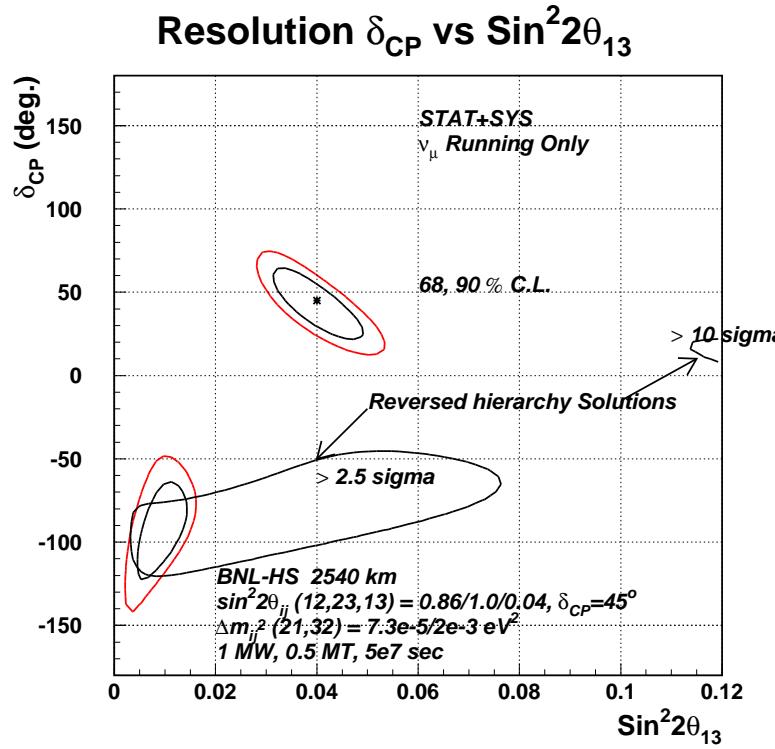
Running for $\sin^2 2\theta_{13}$



$$\Delta m_{32}^2 = 0.002 \text{ eV}^2, \sin^2 2\theta_{13} = 0.04.$$

Reversed Mass Hierarchy
Matter effects included.

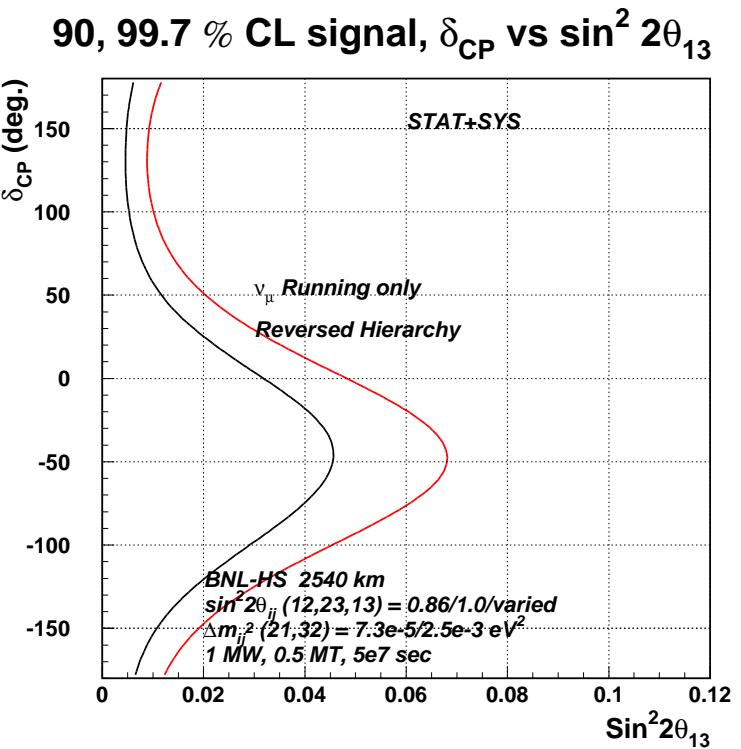
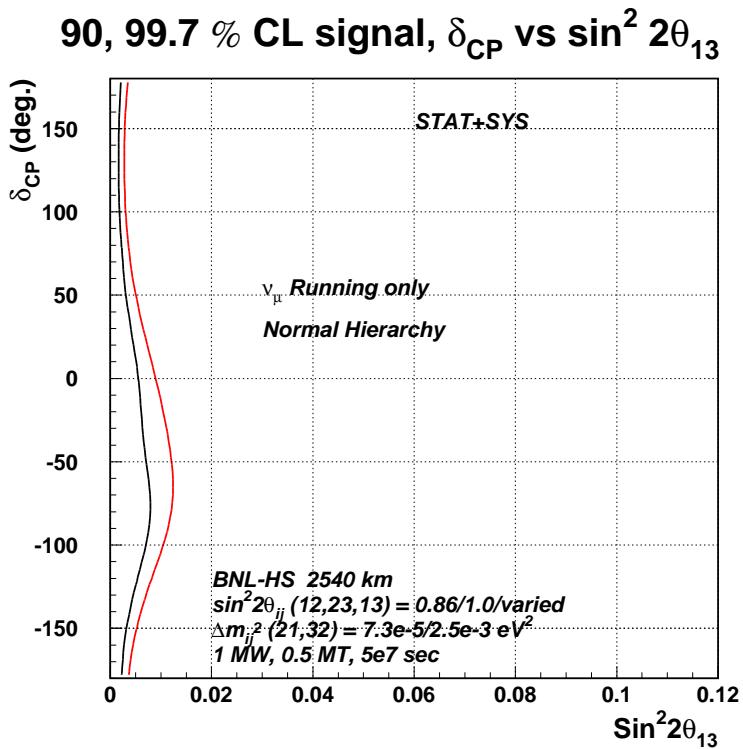
Mass hierarchy after neutrino running



Left: Regular mass hierarchy Right: reversed mass hierarchy.

Mass hierarchy is resolved to 3 sigma with only neutrino running in large part of the parameter space.

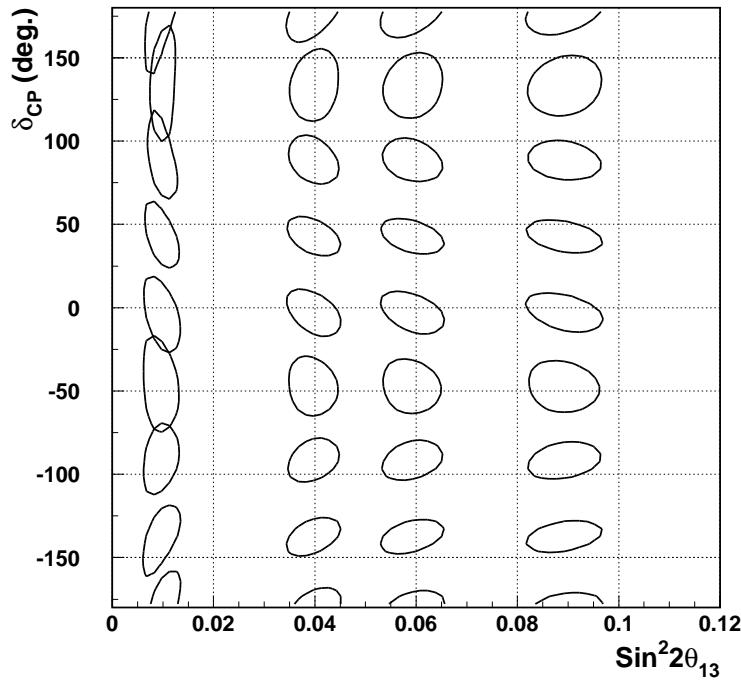
$\sin^2 2\theta_{13}$ sensitivity



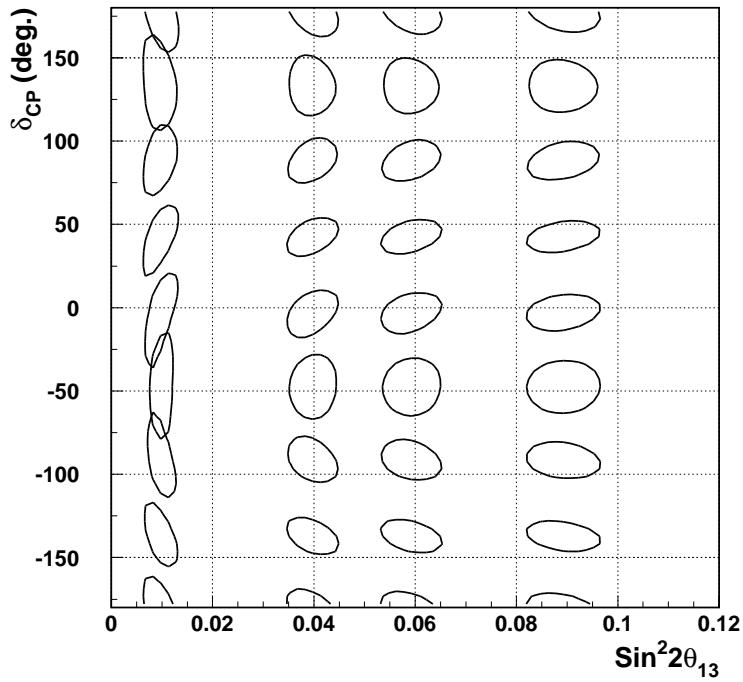
If reversed hierarchy and in the unlucky region, need to run anti-neutrinos.

CP measurement after nu and anti-nu

Regular hierarchy $\nu\nu$ and $\bar{\nu}\bar{\nu}$ running



Reversed hierarchy $\nu\nu$ and $\bar{\nu}\bar{\nu}$ running



Left: Regular mass hierarchy Right: reversed mass hierarchy.

Only the θ_{23} ambiguity is left.
more about this later...

Scnenario 2. Background is worse

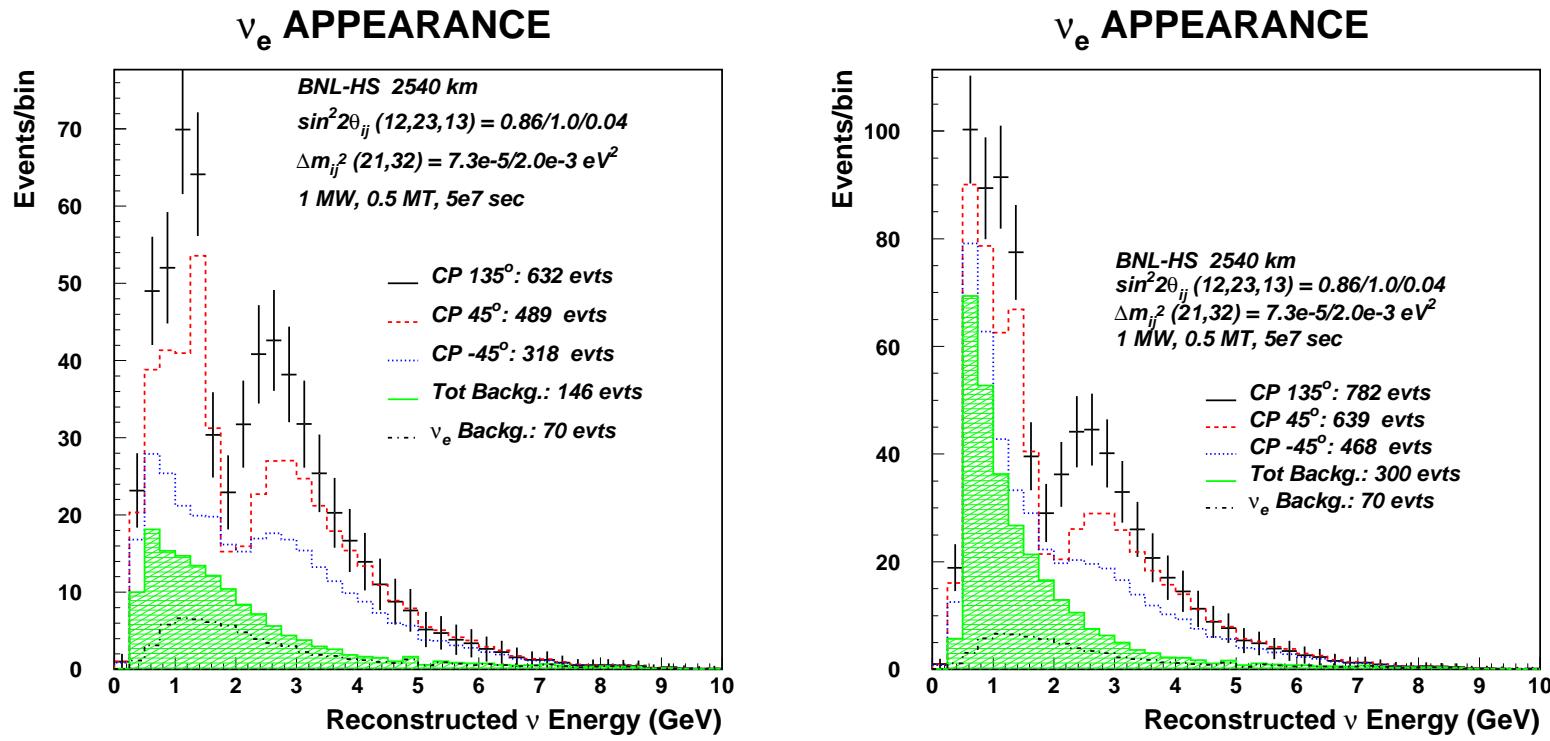
Same running conditions.

Particles within 20 deg. separation cannot be resolved.

- Some effect on disappearance.
- Some effect on appearance with $E > 2.5GeV$.
- Large effect at lower energies.

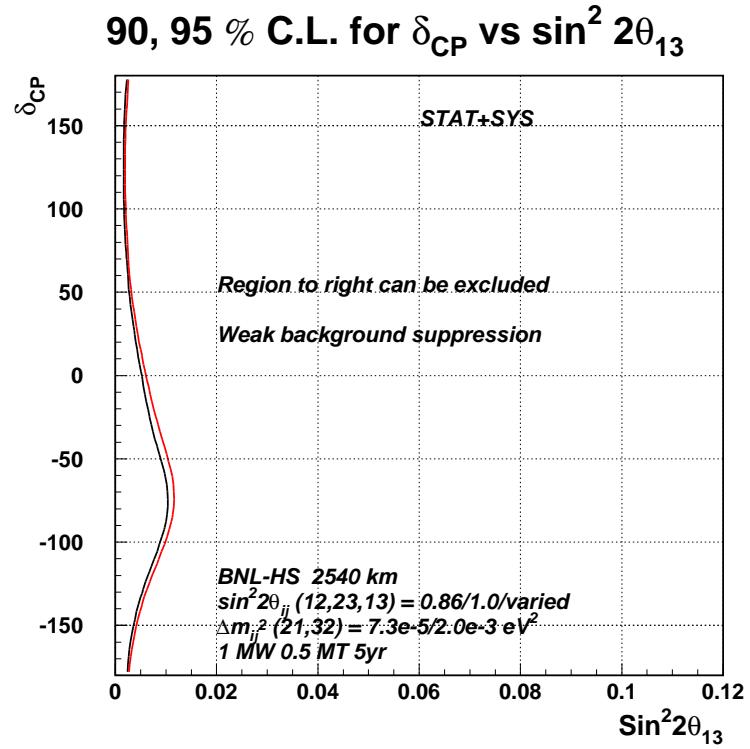
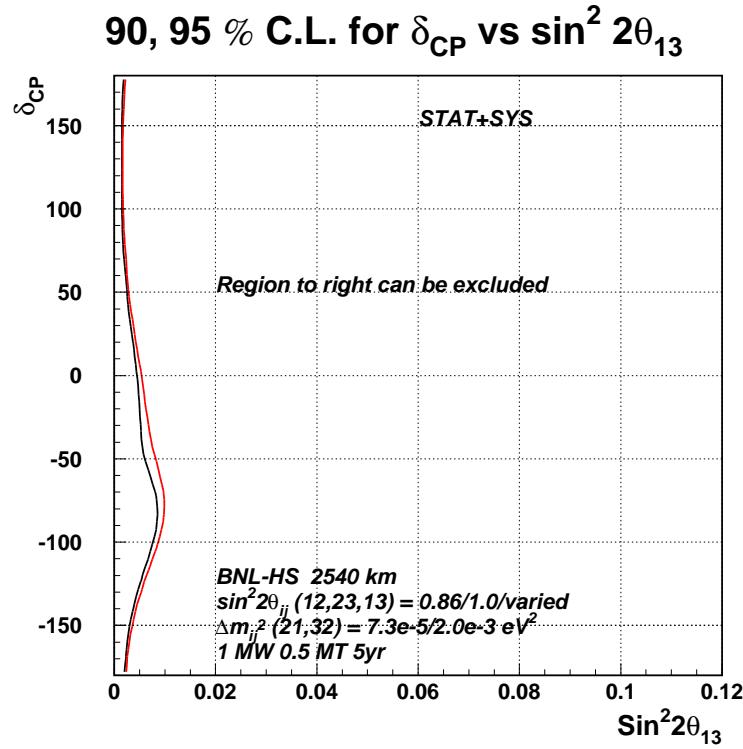
Can background above 2-3 GeV be kept under control by shaping the beam spectrum ?

Worse background spectra



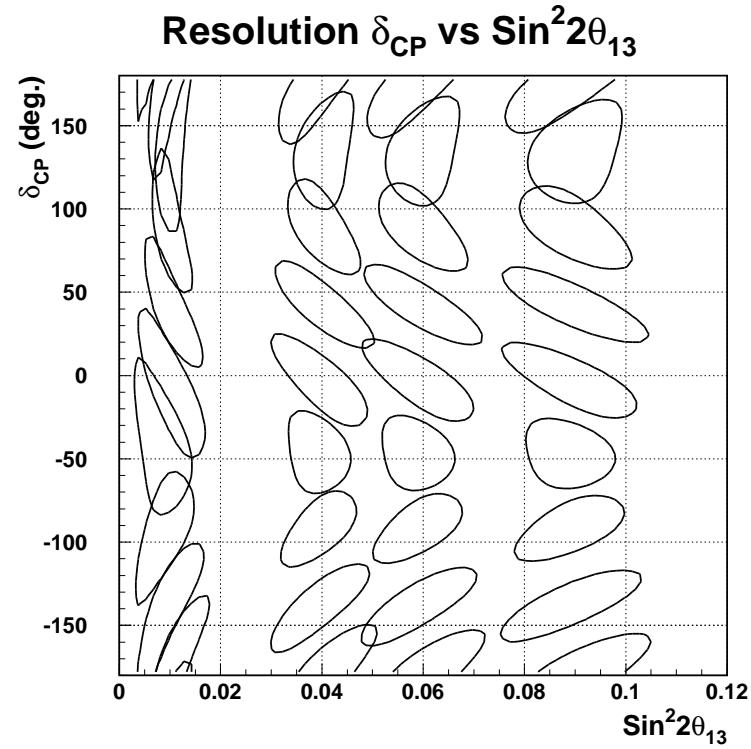
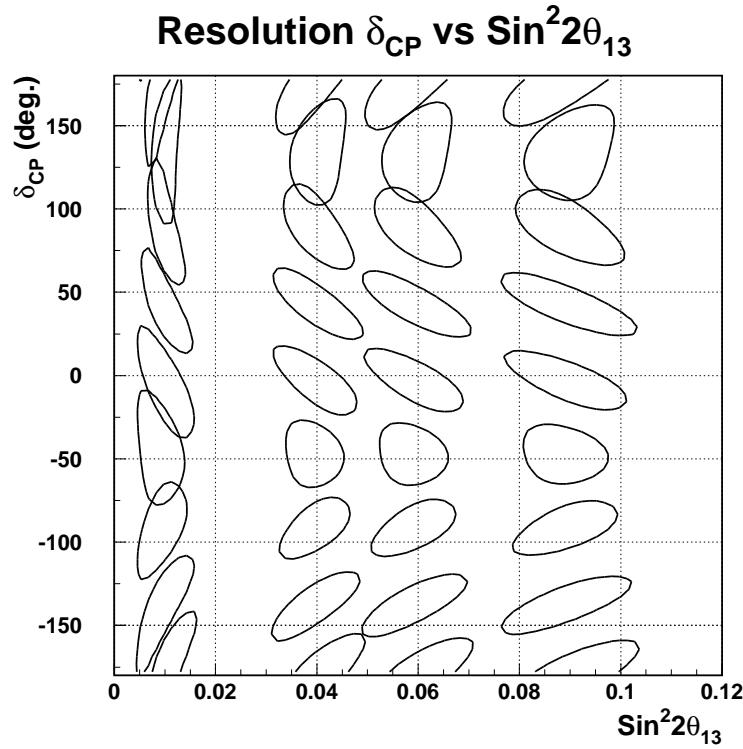
Left: strong rejection Right: weak rejection

Worse background limit



Left: strong rejection Right:weak rejection Limit affected by < 2 :
 Using knowledge of δm_{32}^2 in extracting limit
 Background is confined to lower node.

Worse background CP-measurement



Left: strong rejection Right: weak rejection

Worse background: δ_{CP} worse by $\sim 50\%$

Could one run off-axis to do better ?

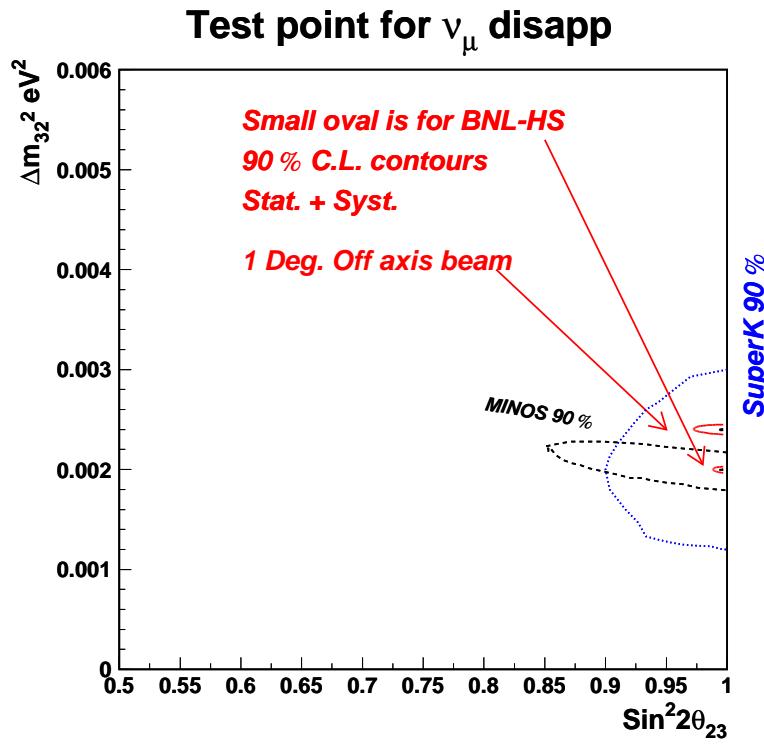
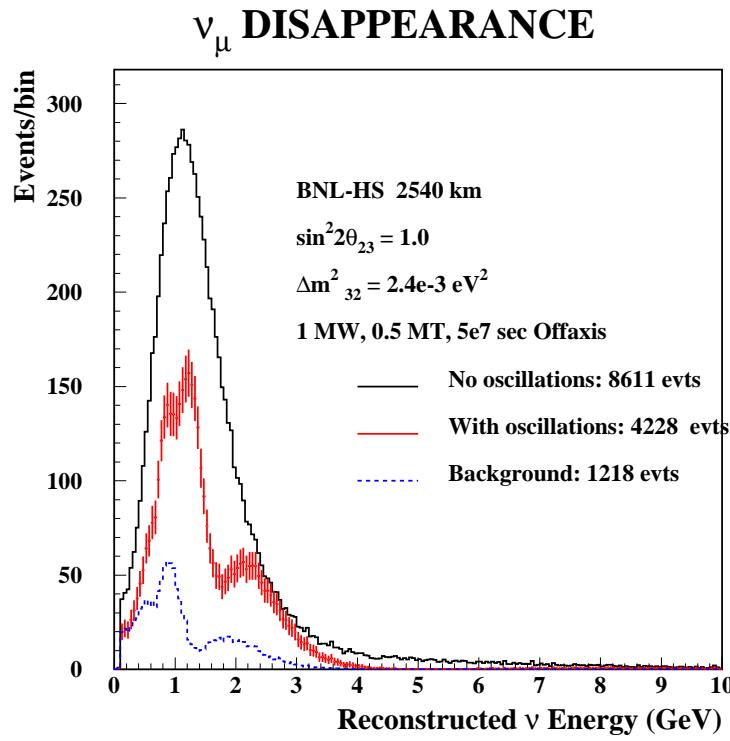
1 Deg. neutrino running

- Better backgrounds for ν_e detection.
- Refine CP measurement.
- Better measurement of θ_{23} and Solar parameters ?

1 MW, 500 kT, Neutrino running, 5×10^7 sec, 2540 km (1.22×10^{22} Protons at 28 GeV.)

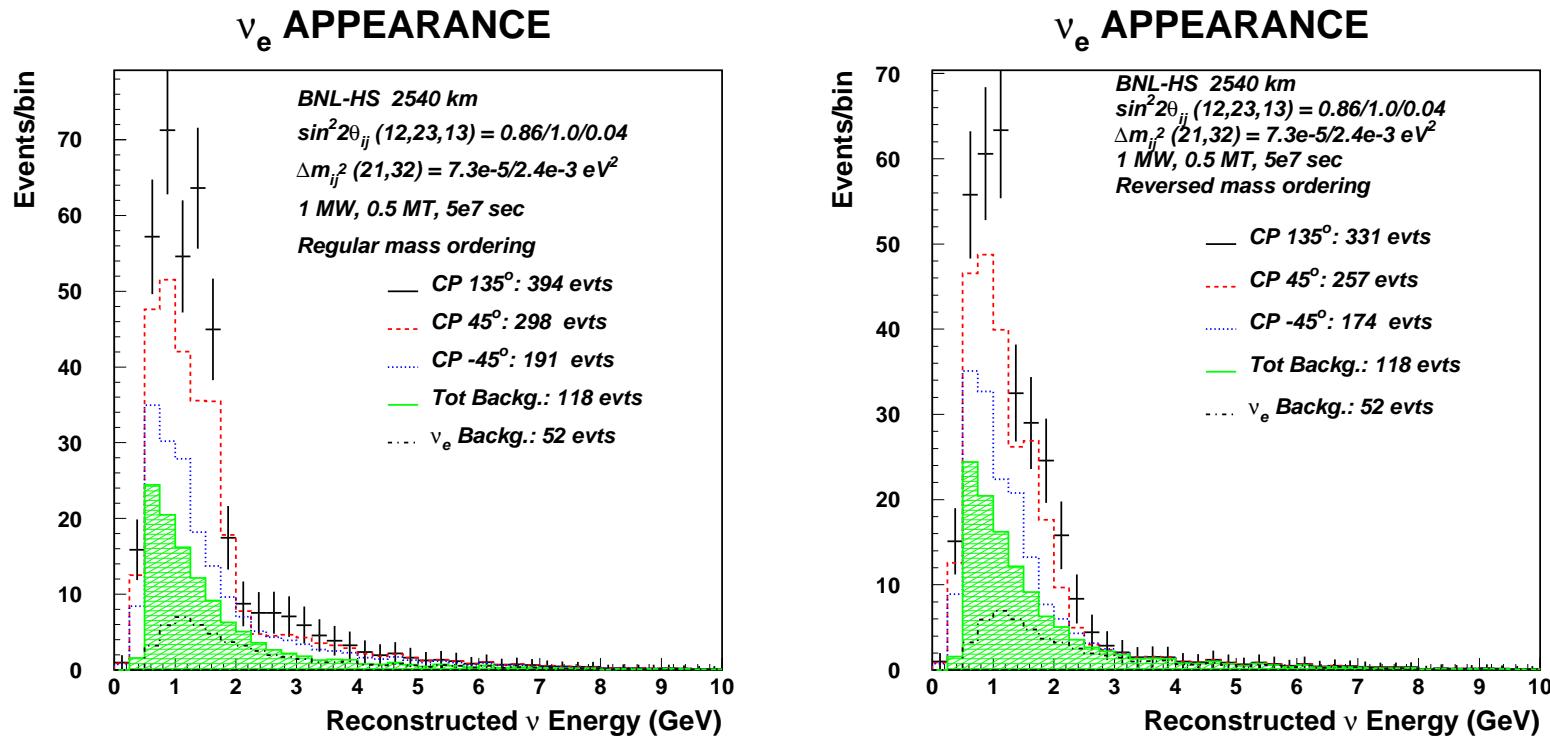
	On axis	1 Deg. Off axis
CC/NC $\nu_\mu N \rightarrow lX$	51800/18323	18931/7081
CC $\nu_e N \rightarrow e^- X$	380	265
QE $\nu_\mu n \rightarrow \mu^- p$	11767	6462
QE $\nu_e n \rightarrow e^- p$	84	69
CC/NC Single π	22053/7741	8445/2996
CC/NC Two π	10143/3557	2394/814
CC $\nu_\tau N \rightarrow \tau^- X$	~ 110	50

1 deg. disappearance



Loss of one node, Reduction in $\sin^2 2\theta_{23}$ resolution.
About factor of 3 compared to wide-band running.

1 deg. appearance



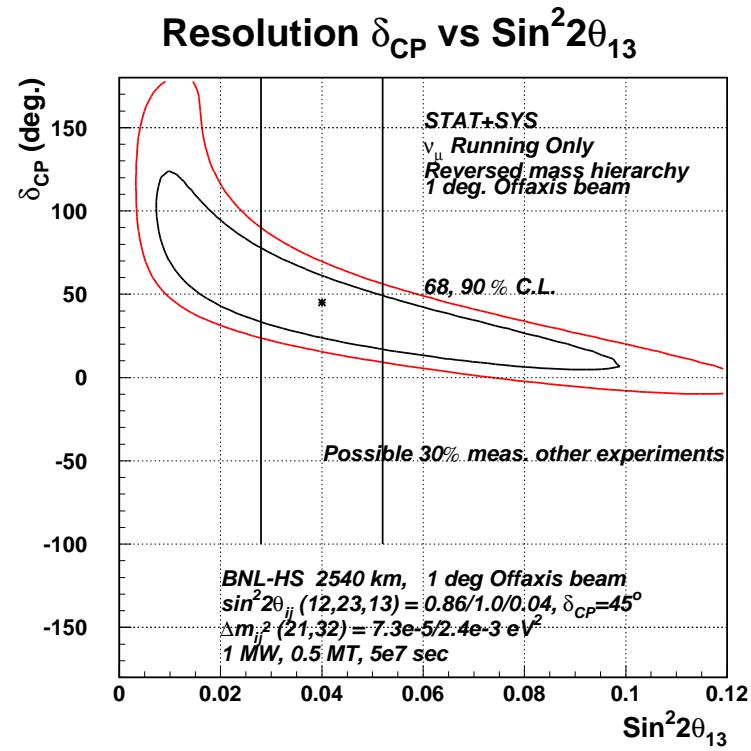
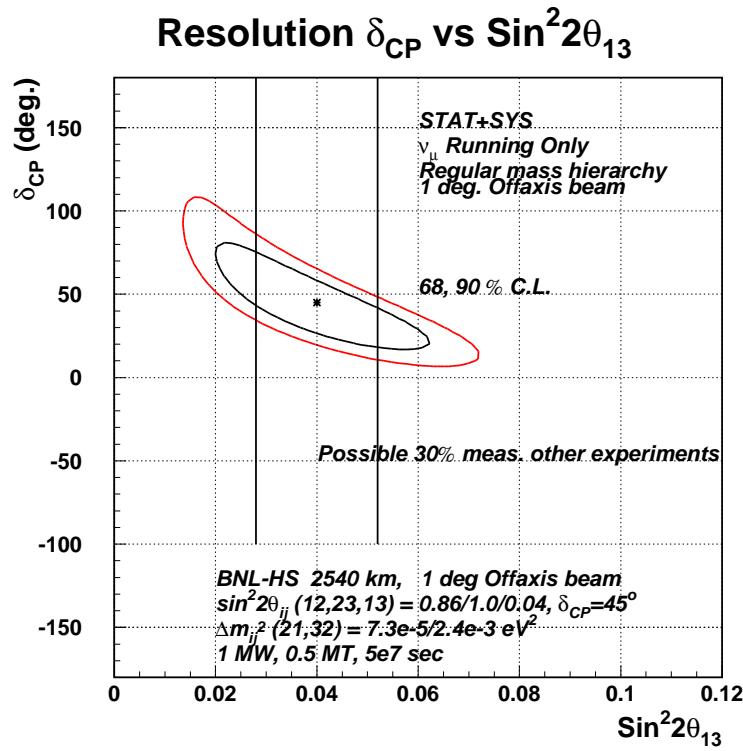
Left: regular mass ordering; Right: reversed mass ordering

Assume weak background rejection

Obtain factor of ~ 3.5 in NC backg. due to 1 deg.

Mass ordering discrimination worse.

1 deg. CP measurement

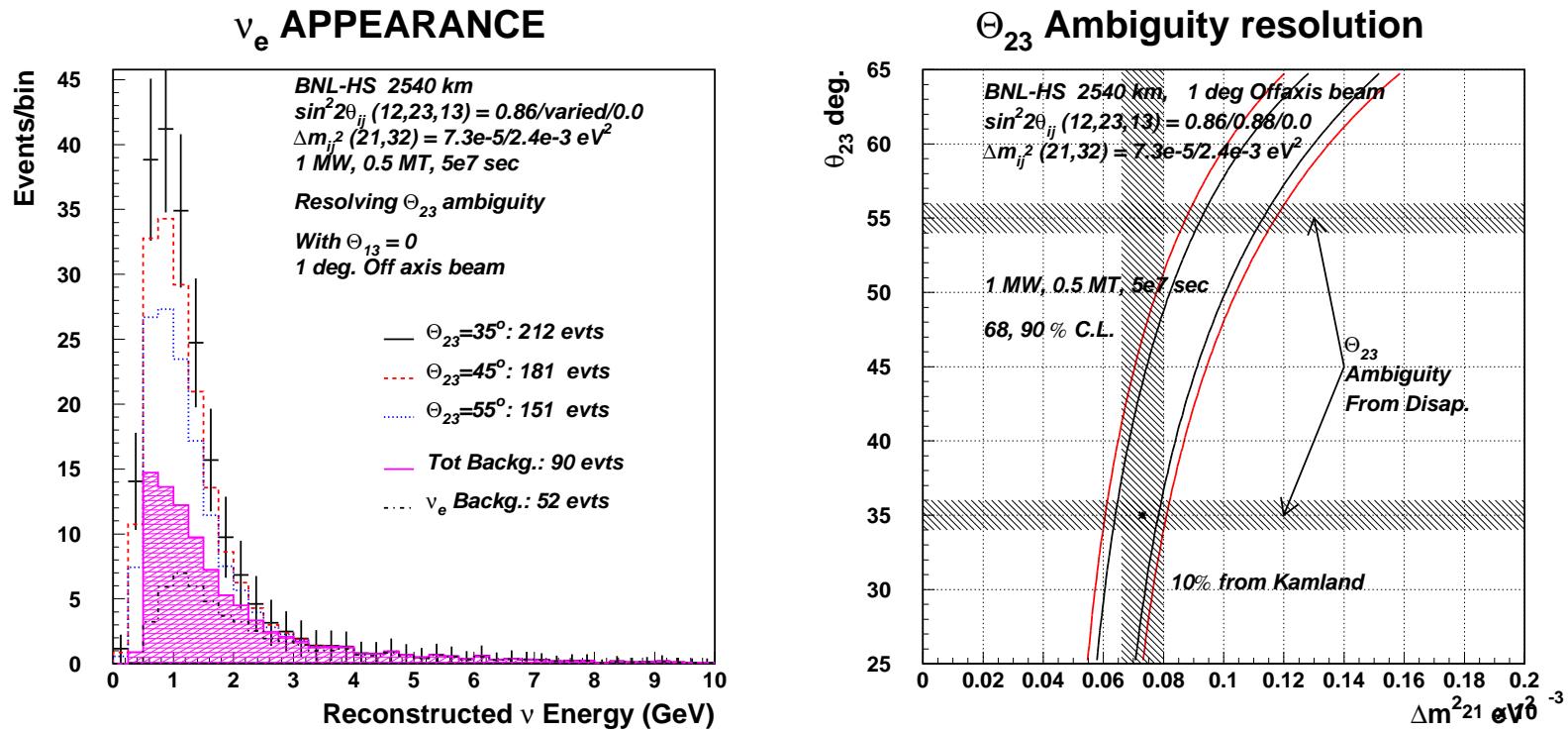


Left: regular mass ordering; Right: reversed mass ordering

Large CP vs. $\sin^2 2\theta_{13}$ correlation.

Will need separate measurement of $\sin^2 2\theta_{13}$.

1 deg. θ_{23} measurement



If $\sin^2 2\theta_{23} < 1.0$ then $\theta_{23} \rightarrow \pi/2 - \theta_{23}$ ambiguity.

Can be resolved with $\nu_\mu \rightarrow \nu_e$ at low energies.

Need strong backg. rej. and off-axis beam.

Use KAMLAND $\sin 2\theta_{12} \times \Delta m_{21}^2$ measurement.

On axis beam to LAR

1 MW, Neutrino running, $5 \times 10^7 sec$

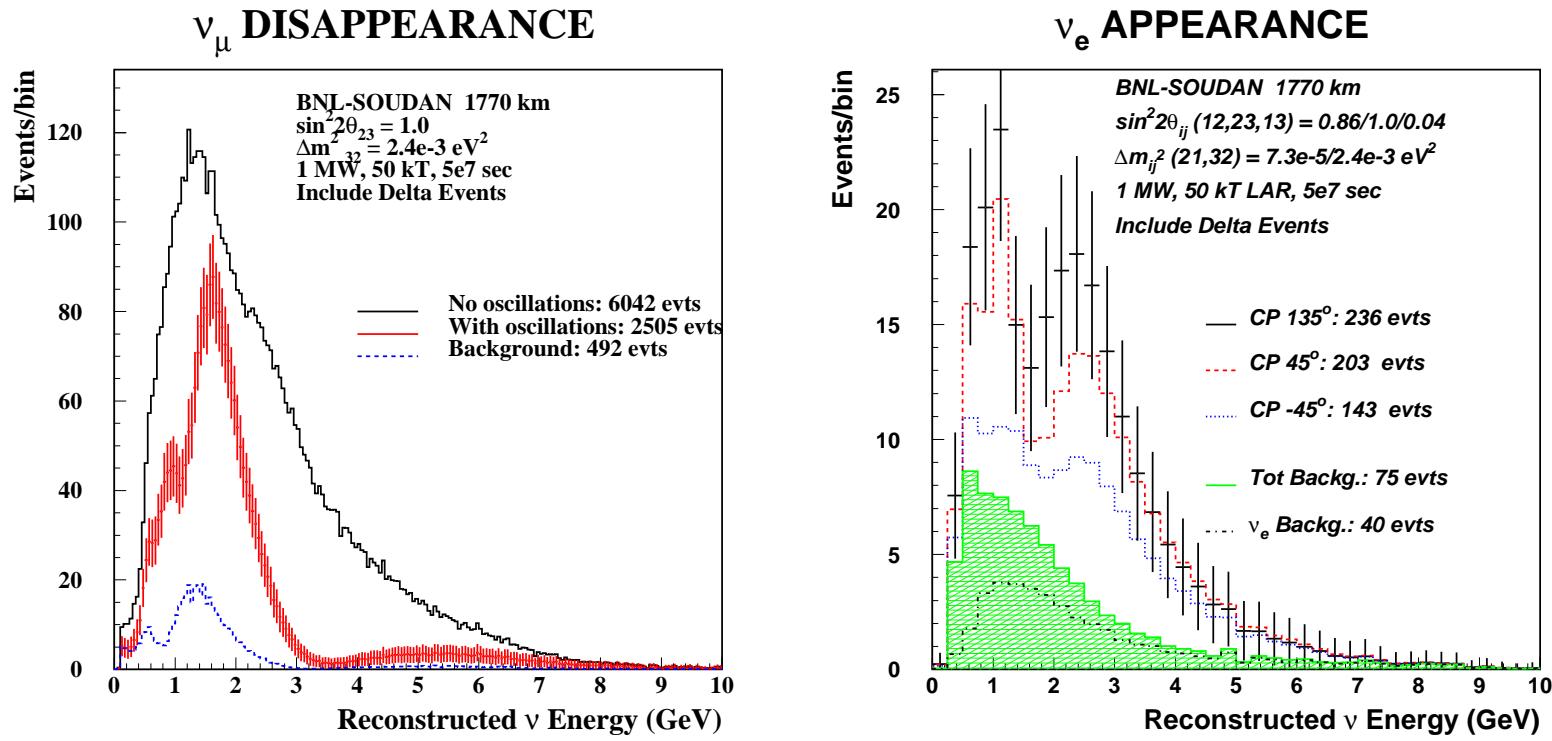
	0.5 MT 2540km H2O	50 kT 1770km LAR
CC/NC $\nu_\mu N \rightarrow l X$	51800/18323	12371/4240
CC $\nu_e N \rightarrow e^- X$	380	93
QE $\nu_\mu n \rightarrow \mu^- p$	11767	3372
QE $\nu_e n \rightarrow e^- p$	84	24
CC/NC Single π	22053/7741	5082/1783
CC/NC Two π	10143/3557	2535/825
CC $\nu_\tau N \rightarrow \tau^- X$	~ 110	25

(Used slightly different spectra for two simulations)

Most physics preserved if event rate can be obtained.

Use all CC ? Go to 2 MW ?

LAR1770 disappearance, appearance



Only 2 nodes visible. Increase yield with delta production events.
Treat all $\nu_l p \rightarrow l^- \pi^+ n$ as if $\nu_l p \rightarrow \Delta^+ l^-$
Reduces resolution below 2 GeV.

Detector Requirements

- Fiducial Mass:
 - $> 500 \text{ kT}$ if using only “clean” events.
Also needed for proton decay and neutrino astrophysics.
 - $\sim 100 \text{ kT}$ if fine grain and use all CC events.
Selected proton decay modes still at the frontier.
- Threshold: $\sim 10 \text{ MeV}$, Dynamic range: contain $5 \text{ GeV } \mu$
- Time res: few ns. Energy resolution: $\sim 10 \%$
- muon/electron discrimination: $< 1 \%$
- Pattern recognition:
 - 1, 2, 3 track separation
 - showering vs. multitrack separation.
 - need factor of 20-30 rejection capability around 1-2 GeV.
- Cost: 300 M\\$ – 1000 M\$

Some comments on detector

- Need much more manpower for detector studies.
- Water Cherenkov
 - 50 kT SuperK is existence proof.
 - Background rejection ? (Yanagisawa)
need another $\times 3 \rightarrow 5$
 - Additional imaging capability (Viren)
- Liquid Argon TPC
 - scale up to 100 kT module ?
 - Current size 300Ton.
 - Needs detailed simulations.
- Any other technology ?

Summary of studies

- Background:
 - Kinematic suppression of backg. above 2 GeV needs to be viable.
 - Could tolerate more background at lower energies by switching to 1 deg. beam if needed.
- 1 Deg. running:
 - Use to compliment the wide band running for CP.
 - Or to find $\nu_\mu \rightarrow \nu_e$ at even larger L/E
 - Could resolve θ_{23} ambiguity if needed.
- Liquid Argon:
 - Need at least 50 kT if large fraction of CC events can be used.
 - Could be closer (1720 km ?) and perform adequately for CP.

Conclusion

A next generation large detector at a national underground facility with a powerful neutrino beam and very long baseline.
Unique, and broad range of physics.

- Known Knowns: Δm_{32}^2 , $\sin^2 2\theta_{23}$, measure better.
 Δm_{21}^2 $\sin^2 2\theta_{12}$ constrain with appearance.
- Known Unknowns: θ_{13} , δ_{CP} , θ_{23} ambiguity.
sign of Δm_{32}^2 and matter effect
- Unknown Unknowns: Over-constrain the 3-generation mixing picture.